

We contradict binary nature of LQ And. Radial velocities of H α emission are bound to changes less than 2 km/s. We were not able to detect any periodic variability in radial velocities of core and wings of H α emission nor of line He I 6678 Å. We discuss effect of telluric line movement on precision of our measurement. Detected period of 0,415 d in intensity ratio I_V/I_R needs additional confirmation. The star overcame striking increase in intensity of the emission accompanied by increase in overall brightness and shortening of photometric period from 0,3094 d to 0,3083 d. Our estimated interval of rotational periods of the star lies in between this photometric period and double of it, which shows that rotation might not be responsible for the photometric changes.